# Search for radiative $D_s$ decays

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Abstract. The study of weak radiative decays of charmed mesons is still in its developing stage. The weak decays of D mesons pose challenges due to significant final-state interactions. However, decays mediated by  $c \rightarrow u\gamma$  transitions can be affected by potential contributions coming from the non-minimal supersymmetry, which is an new physics scenario. The ratio of branching fractions for radiative  $D^0$  decays could be violated already in the SM framework, while a similar ratio for  $D_s^+$  radiative decays offers much better prospects for new physics. We present herein the first sensitivity study of the radiative charm decays  $D_s^+ \rightarrow \rho^+ \gamma$  and  $D_s^+ \rightarrow K^{*+}\gamma$  with data collected by the Belle experiment.

Keywords: radiative, charm decays, Belle

#### 1 Introduction

In the Standard Model (SM), the physics of charmed mesons faces certain challenges compared to strange and beauty mesons because the CP asymmetries and  $D^0 - \overline{D}^0$  oscillations are small. Further, the weak decays of D mesons are difficult to investigate due to significant final-state interactions. However, it has been pointed out that the oscillations and  $c \to u\gamma$  decays might have some contributions coming from the non-minimal supersymmetry (an NP scenario). Therefore, one can search for NP using  $c \to u\gamma$  transitions. It was suggested that the NP would result in a deviation from the ratio of branching fractions [1]:

$$R_{\rho/\omega} \equiv \frac{\Gamma(D^0 \to \rho^0/\omega\gamma)}{\Gamma(D^0 \to \bar{K}^{*0}\gamma)} = \frac{\tan^2\theta_c}{2},\tag{1}$$

where  $\theta_c$  is the Cabibbo angle. In order to find the best mode to study  $c \to u\gamma$  transitions, the ratios between various Cabibbo-suppressed and Cabibbo-allowed radiative decays of charmed mesons are calculated within the SM. It has been noticed that Eq. (1) could be violated already in the SM framework because of a large, unknown correction, while a similar ratio for  $D_s^+$  radiative decays

$$R_K \equiv \frac{\Gamma(D_s^+ \to K^{*+}\gamma)}{\Gamma(D_s^+ \to \rho^+\gamma)} = \tan^2\theta_c, \tag{2}$$

offers a much better probe for an NP signal, as the latter is less sensitive to the SM correction [2]. Long-distance(LD), non-perturbative processes dominate these decays, potentially enhancing Branching Fractions(BFs), basically 2 N. S. Ipsita et al.

to test the Quantum Chromodynamics(QCD) based calculations of LD dynamics. The BF of  $D_s^+ \rightarrow \rho^+ \gamma \ [D_s^+ \rightarrow K^{*+} \gamma]$  mode is expected to lie within the range of  $O(10^{-5}) - O(10^{-3}) \ [O(10^{-8}) - O(10^{-4})]$ , according to the predictions of different theoretical models [7], which are quite divergent for the  $D_s$  decay mode. The BF of  $D_s^+ \rightarrow \gamma \rho(770)^+$  is measured to be  $(2.2 \pm 0.9 \pm 0.2) \times 10^{-4}$ , corresponding to an upper limit of  $6.1 \times 10^{-4}$  at the 90% confidence level by BESIII collaboration [8]. We present herein the first sensitivity study of these radiative  $D_s$  meson decays with data collected by the Belle experiment.

## 2 KEKB and Belle

The Belle detector was located at an interaction point of the KEKB asymmetricenergy  $e^+e^-$  collider. It was a large-solid-angle magnetic spectrometer comprising six subdetectors [3], namely silicon vertex detector, central drift chamber, aerogel Cherenkov counter, time-of-flight counter, CsI(Tl) crystal electromagnetic calorimeter, and  $K_L^0$  and muon detector.

#### 3 Analysis Strategy

We optimize the selection of signal candidates using simulated samples generated with the EvtGen [4] and Geant packages [5]. We reconstruct  $D_s^+$  from  $D_s^+ \to \rho^+ \gamma$  and  $D_s^+ \to K^{*+} \gamma$ , where  $\rho^+ \to \pi^+ \pi^0$ ,  $K^{*+} \to K_S^0 \pi^+$  and  $K^{*+} \to K^+ \pi^0$ . These studies are based on MC samples corresponding to an integrated luminosity of 711 fb<sup>-1</sup>. The kinematic variable that distinguishes signal from background is the invariant mass of  $D_s^+$ . A  $\pi^0$  veto is implemented to suppress the huge background coming from  $\pi^0$  decays. We perform a dedicated background MC study in which the continuum background is found to be dominant. We employ multivariate analysis (MVA) based on the FastBDT package [6] to get rid of uds background, i.e.,  $e^+e^- \to u\bar{u}, d\bar{d}, s\bar{s}$ . After requiring the MVA output to be greater than 0.4 (0.5) for the  $D_s^+ \to \rho^+ \gamma$  ( $D_s^+ \to K^{*+} \gamma$ ) decay mode, we find a rejection of 65% (76%) of uds background at the cost of 10% (24%) of signal loss. The reconstruction efficiency is 0.5\%, 3.1% and 0.8% for  $D_s^+ \to \rho^+ \gamma$  and  $D_s^+ \to K^{*+} \gamma$  modes. For  $D_s^+ \to [\rho^+ \to \pi^+ \pi^0] \gamma$  and  $D_s^+ \to \rho^+ \eta$  and  $D^{*+} \to D^0 \pi^+$ , while for  $D_s^+ \to [K^{*+} \to K_s^0 \pi^+] \gamma$ , it is mostly from  $D^0 \to K_S^0 \pi^0$  and  $D^0 \to K_S^0 \eta$  decay modes.

#### 4 Control Sample Study

We use the peaking backgrounds  $D_s^+ \to \rho^+[\eta \to \gamma\gamma]$ ,  $D^{*0} \to [D^0 \to K_S^0 \eta]\gamma$  and  $D^{*0} \to [D^0 \to K_S^0 \pi^0]\gamma$ , as our control sample to validate the signal extraction procedure and to calibrate possible differences in the signal resolution between data and simulation. Figure 1 shows the unbinned maximum-likelihood fit performed on  $M_{D_s}$  and  $M_{D^0}$  for (a)  $D_s^+ \to \rho^+\eta$ , (b)  $D^0 \to K_S^0 \pi^0$ , and (c)  $D^0 \to K_S^0 \eta^0$ 



**Fig. 1.** Fitted distribution of Invariant mass of  $D^0$  and  $D_s^+$  for (a)  $D^0 \to K_S^0 \pi^0$ , (b)  $D^0 \to K_S^0 \eta$  and (c)  $D_s^+ \to \rho^+ \eta$  decay modes.

decay modes using MC samples corresponding to an integrated luminosity of 711  ${\rm fb}^{-1}.$ 

### 5 Signal Extraction

We have performed 1D unbinned maximum likelihood fit to extract signal yield. Figure 2 shows the total fitted distribution of  $M_{D_s}$  for (a)  $D_s^+ \to \rho^+ \gamma$ , (b)  $D_s^+ \to [K^{*+} \to K_s^0 \pi^+] \gamma$  and (c)  $D_s^+ \to [K^{*+} \to K^+ \pi^0] \gamma$  decay modes, respectively.

# 6 Preliminary Results and Outlook

We are expecting 150-200, 20-30 and 8-12 events for  $D_s^+ \to \rho^+ \gamma$ ,  $D_s^+ \to [K^{*+} \to K_s^0 \pi^+] \gamma$  and  $D_s^+ \to [K^{*+} \to K^+ \pi^0] \gamma$  decay modes, assuming a branching fraction of  $10^{-4}(10^{-5})$ , with a data sample corresponding to an integrated luminosity of 921 fb<sup>-1</sup>.



**Fig. 2.** Fitted distribution of Invariant mass of  $D_s^+$  for  $(a)D_s^+ \to \rho^+\gamma$ ,  $(b)D_s^+ \to [K^{*+} \to K_s^0\pi^+]\gamma$  and  $(c)D_s^+ \to [K^{*+} \to K^+\pi^0]\gamma$  decay modes.

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